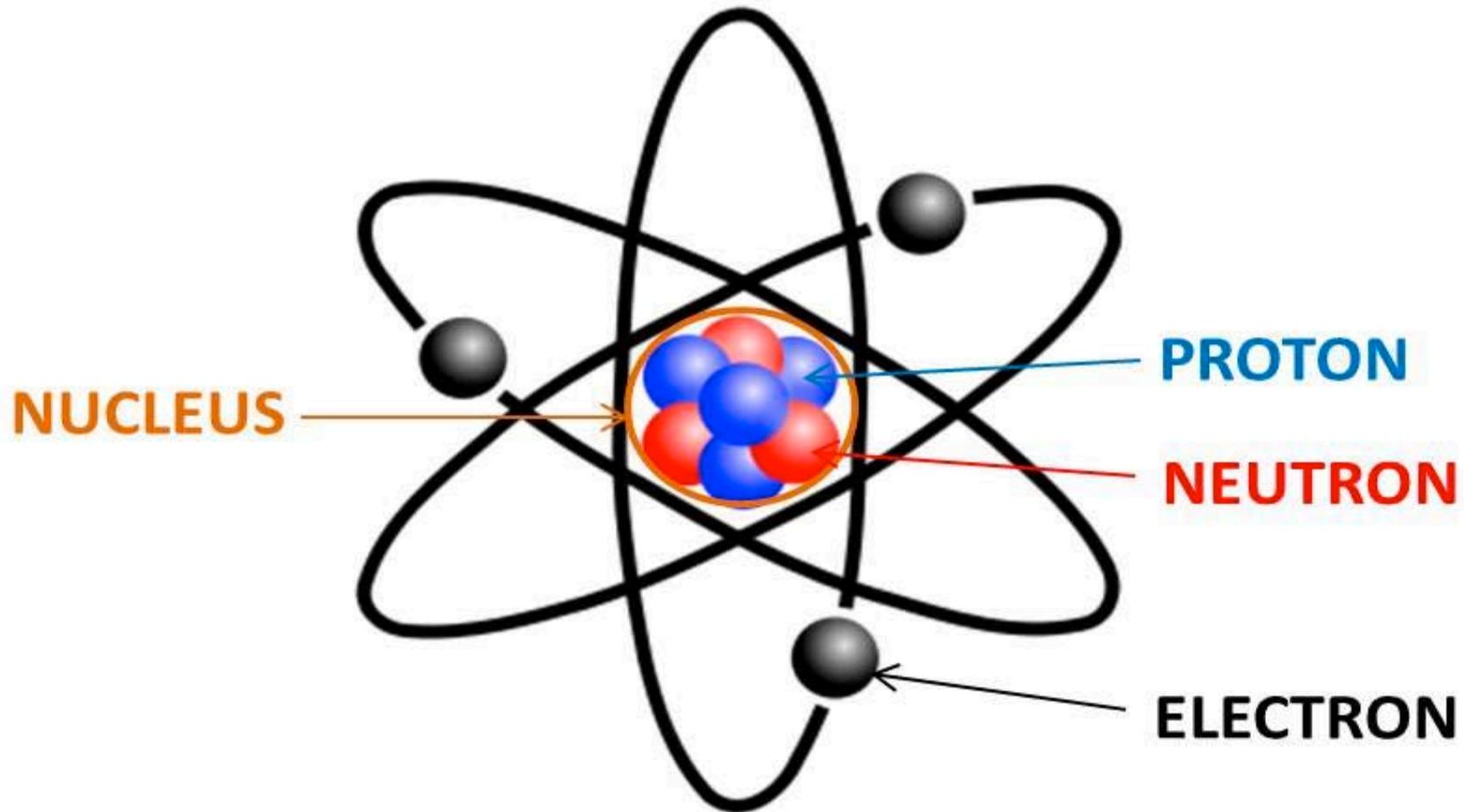
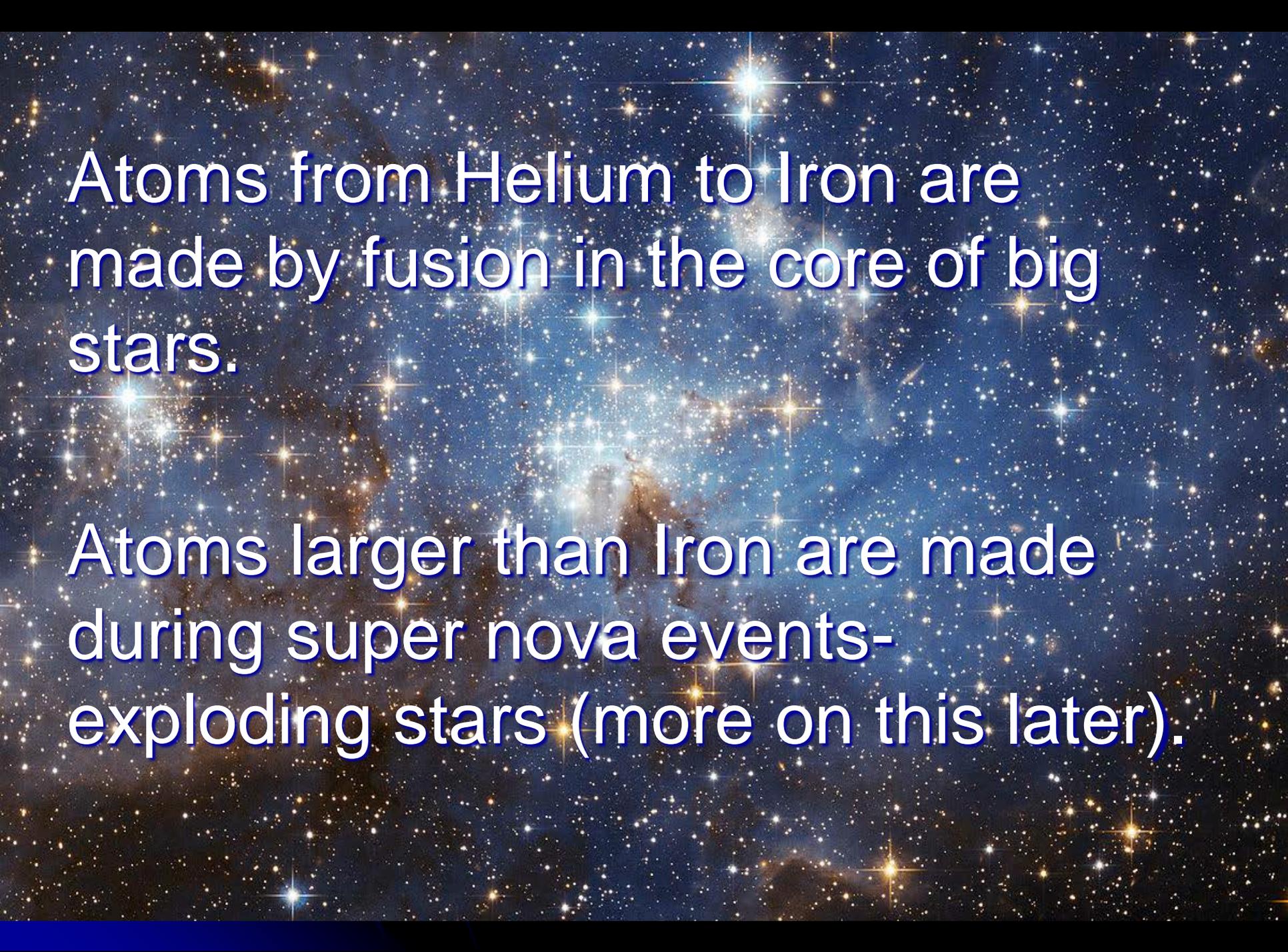


Where did atoms come from?

The first atoms of hydrogen and helium were created during the big bang.



A background image of a starry night sky. The sky is dark blue and black, filled with numerous bright stars of various colors, including white, yellow, and orange. In the center, there is a prominent nebula with a mix of blue and white colors, surrounded by a faint, glowing ring of light. The overall scene is a rich, multi-colored star field.

Atoms from Helium to Iron are made by fusion in the core of big stars.

Atoms larger than Iron are made during super nova events- exploding stars (more on this later).

IUPAC Periodic Table of the Elements

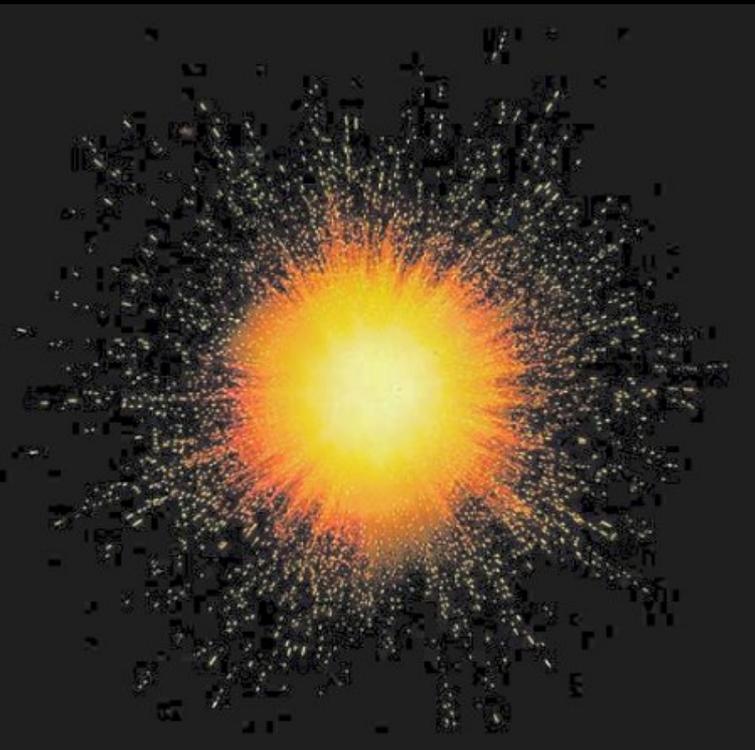
1 H hydrogen 1.008 [1.0078, 1.0082]																	18 He helium 4.0026
3 Li lithium 6.94 [6.938, 6.997]	4 Be beryllium 9.0122											13 B boron 10.81 [10.806, 10.821]	14 C carbon 12.011 [12.009, 12.012]	15 N nitrogen 14.007 [14.006, 14.008]	16 O oxygen 15.999 [15.999, 16.000]	17 F fluorine 18.998	10 Ne neon 20.180
11 Na sodium 22.990	12 Mg magnesium 24.305 [24.304, 24.307]											13 Al aluminium 26.982	14 Si silicon 28.085 [28.084, 28.086]	15 P phosphorus 30.974	16 S sulfur 32.06 [32.059, 32.076]	17 Cl chlorine 35.45 [35.446, 35.457]	18 Ar argon 39.948
19 K potassium 39.098	20 Ca calcium 40.078(4)	21 Sc scandium 44.956	22 Ti titanium 47.867	23 V vanadium 50.942	24 Cr chromium 51.996	25 Mn manganese 54.938	26 Fe iron 55.845(2)	27 Co cobalt 58.933	28 Ni nickel 58.693	29 Cu copper 63.546(3)	30 Zn zinc 65.38(2)	31 Ga gallium 69.723	32 Ge germanium 72.630(8)	33 As arsenic 74.922	34 Se selenium 78.971(8)	35 Br bromine 79.904 [79.901, 79.907]	36 Kr krypton 83.798(2)
37 Rb rubidium 85.468	38 Sr strontium 87.62	39 Y yttrium 88.906	40 Zr zirconium 91.224(2)	41 Nb niobium 92.906	42 Mo molybdenum 95.95	43 Tc technetium	44 Ru ruthenium 101.07(2)	45 Rh rhodium 102.91	46 Pd palladium 106.42	47 Ag silver 107.87	48 Cd cadmium 112.41	49 In indium 114.82	50 Sn tin 118.71	51 Sb antimony 121.76	52 Te tellurium 127.60(3)	53 I iodine 126.90	54 Xe xenon 131.29
55 Cs caesium 132.91	56 Ba barium 137.33	57-71 lanthanoids	72 Hf hafnium 178.49(2)	73 Ta tantalum 180.95	74 W tungsten 183.84	75 Re rhenium 186.21	76 Os osmium 190.23(3)	77 Ir iridium 192.22	78 Pt platinum 195.08	79 Au gold 196.97	80 Hg mercury 200.59	81 Tl thallium 204.38 [204.38, 204.39]	82 Pb lead 207.2	83 Bi bismuth 208.98	84 Po polonium	85 At astatine	86 Rn radon
87 Fr francium	88 Ra radium	89-103 actinoids	104 Rf rutherfordium	105 Db dubnium	106 Sg seaborgium	107 Bh bohrium	108 Hs hassium	109 Mt meitnerium	110 Ds darmstadtium	111 Rg roentgenium	112 Cn copernicium	113 Nh nihonium	114 Fl flerovium	115 Mc moscovium	116 Lv livermorium	117 Ts tennessine	118 Og oganesson

57 La lanthanum 138.91	58 Ce cerium 140.12	59 Pr praseodymium 140.91	60 Nd neodymium 144.24	61 Pm promethium	62 Sm samarium 150.36(2)	63 Eu europium 151.96	64 Gd gadolinium 157.25(3)	65 Tb terbium 158.93	66 Dy dysprosium 162.50	67 Ho holmium 164.93	68 Er erbium 167.26	69 Tm thulium 168.93	70 Yb ytterbium 173.05	71 Lu lutetium 174.97
89 Ac actinium 227.03	90 Th thorium 232.04	91 Pa protactinium 231.04	92 U uranium 238.03	93 Np neptunium	94 Pu plutonium	95 Am americium	96 Cm curium	97 Bk berkelium	98 Cf californium	99 Es einsteinium	100 Fm fermium	101 Md mendelevium	102 No nobelium	103 Lr lawrencium



INTERNATIONAL UNION OF
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WHAT EXACTLY IS THE BIG BANG THEORY?



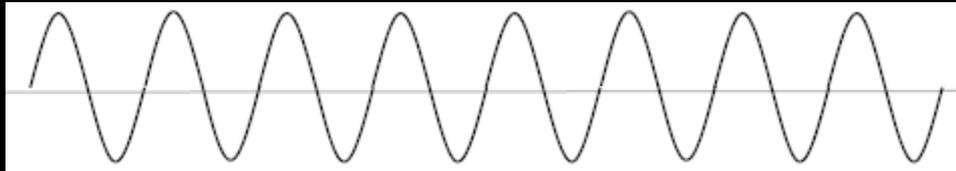
Scientists have calculated that
the universe is approximately
13.7 billion years old.



How do we know this?

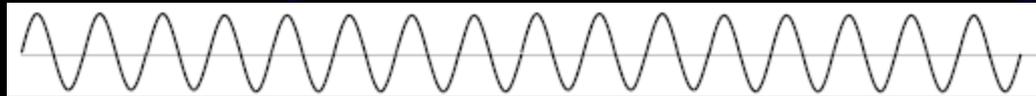
- Currently, most things in the universe are moving apart (expanding) at a constant rate.
- If this expansion was reversed and everything in the universe started to move toward one another (contract) it would take 13.7 billion years for everything to come back together.

How do we know that most objects are moving away from us?



← Normal wave

- When a wave is moving toward you, the wavelengths compress together.

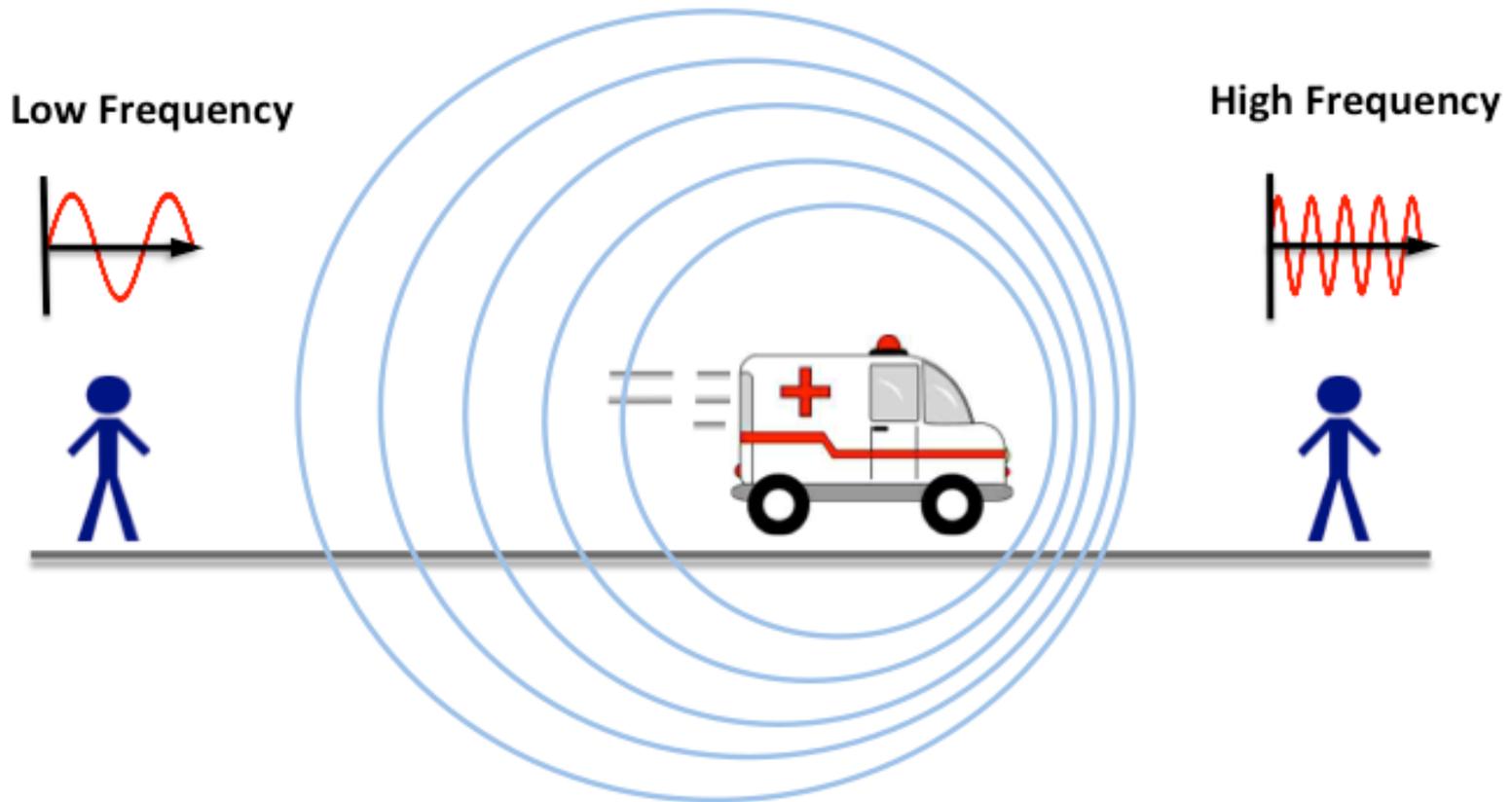


- When a wave is moving away from you the wavelengths expand.



- This happens in light & sound waves

Doppler Effect

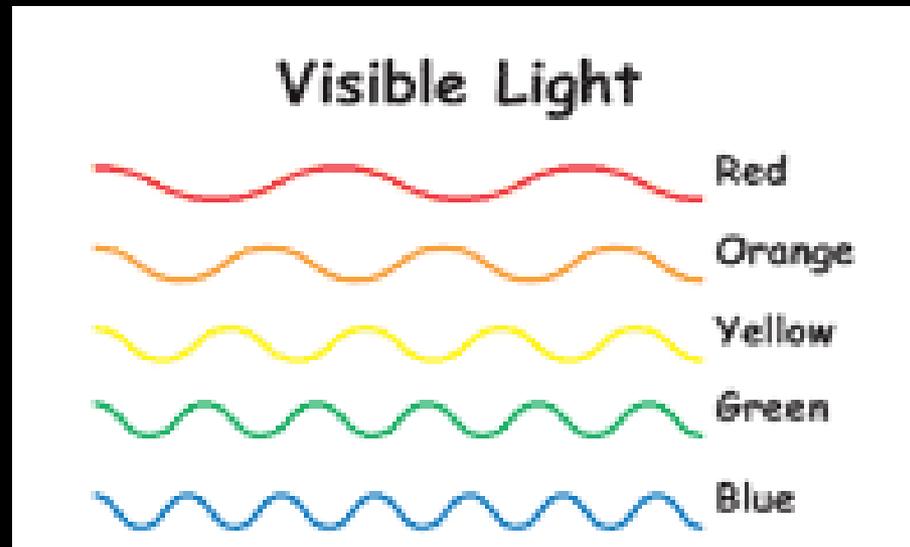


Doppler effect sound wave:

<https://www.youtube.com/watch?v=p-hBCcmCUPg>

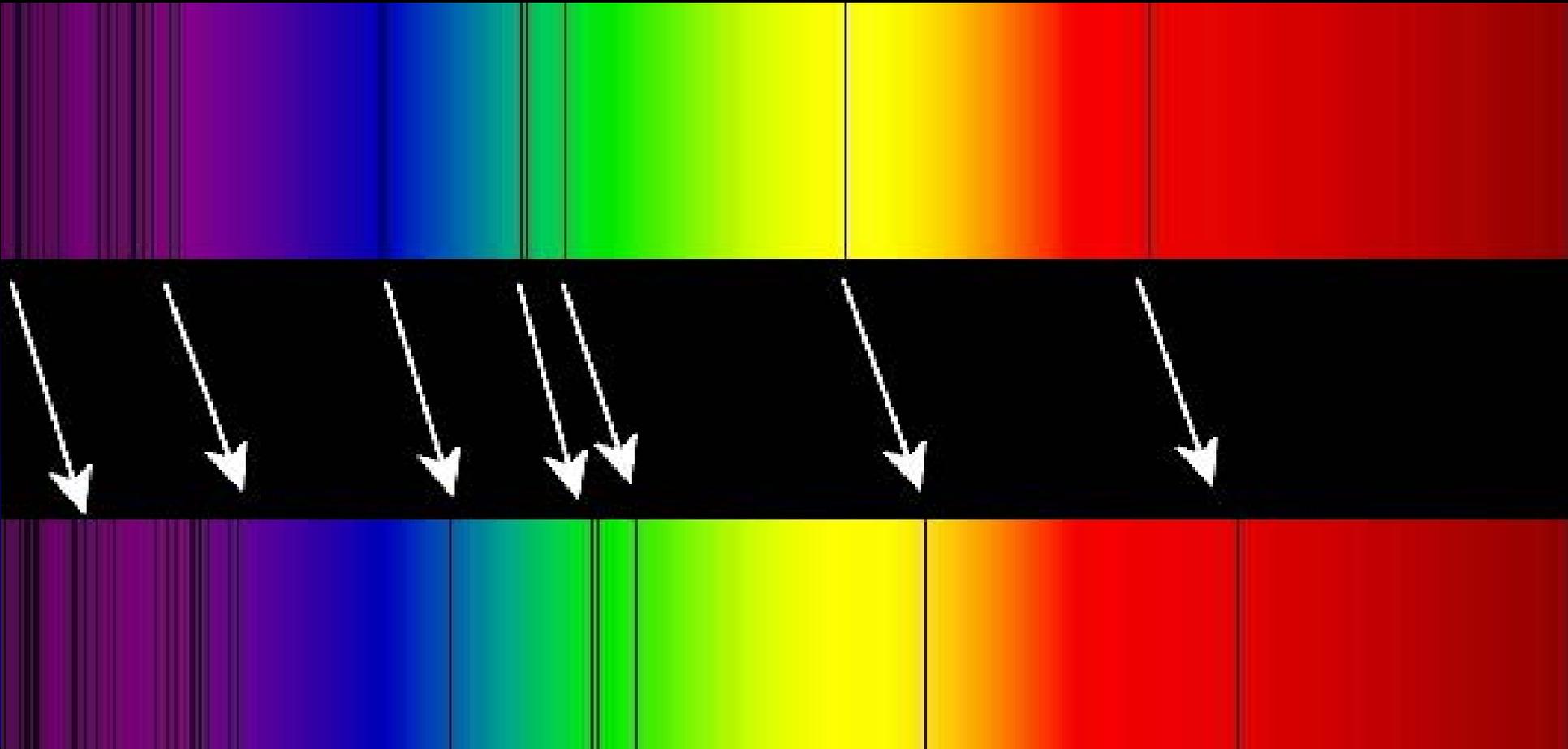
Red Shifts:

- Light coming from most stars appears to be moving away from us.
- When light waves are expanded they change color toward the red end of the color spectra.



Red Shift

A stars spectra lines move into the red end of the color spectrum because the wavelengths are extended.



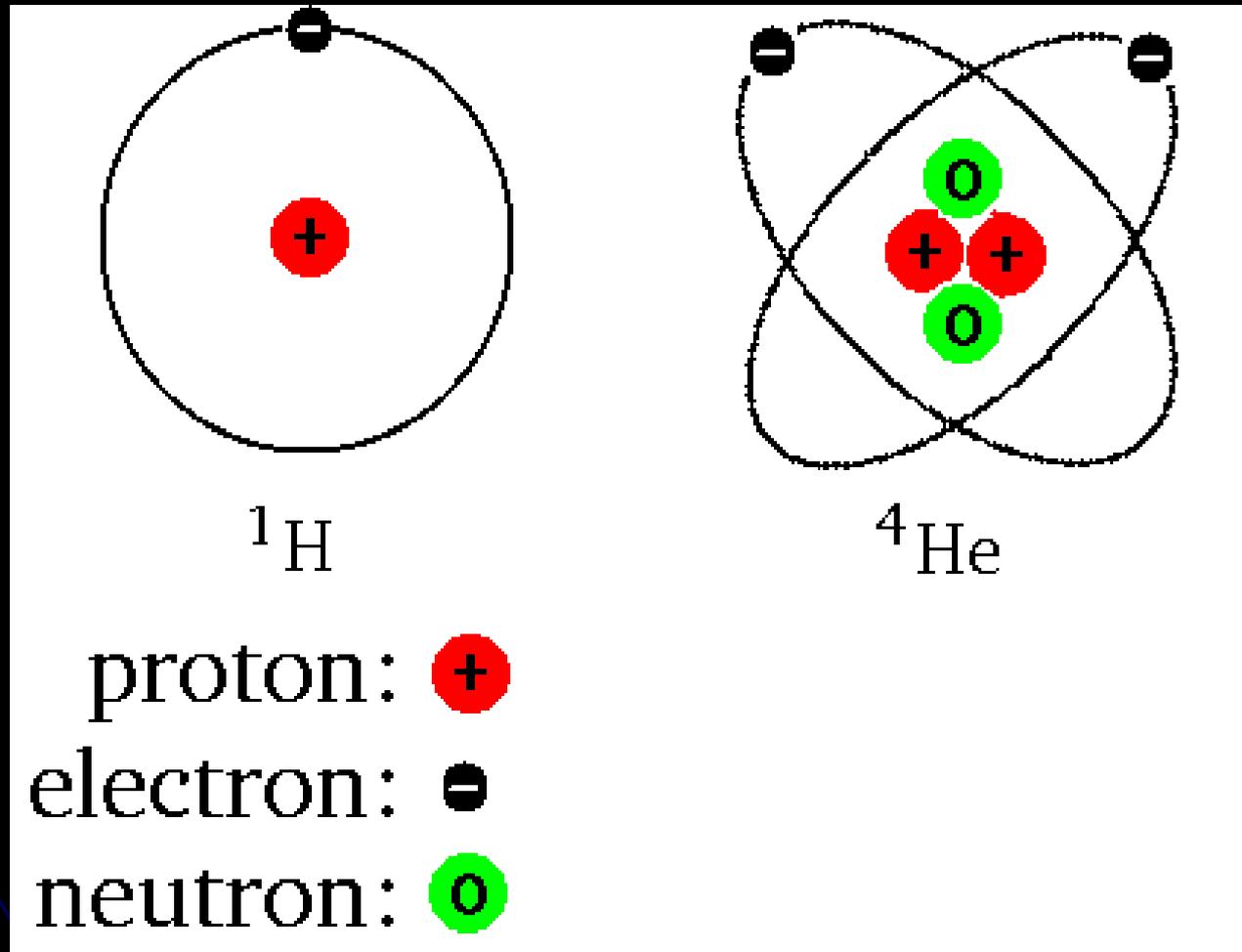


What is the Big Bang?

- **(Step 1)** Approximately 13.7 billion years ago, all the matter & energy in the universe was concentrated into a region smaller than a pen head.
- This region then began to expand & cool at an incredibly rapid rate.

How could the entire universe be squeezed so small?

- Astronomers believe that atoms, as we know them, did not exist.



ATOM:

- The atoms were broken into their smallest pieces- quarks.
- The quarks can be stacked very close together, without all of the empty space that atoms usually have.
- This allowed the entire universe to be squeezed down to the size of a pin head!



What started the expansion?

- (Step 2) The 4 fundamental forces came together in the Unified Force Theory!



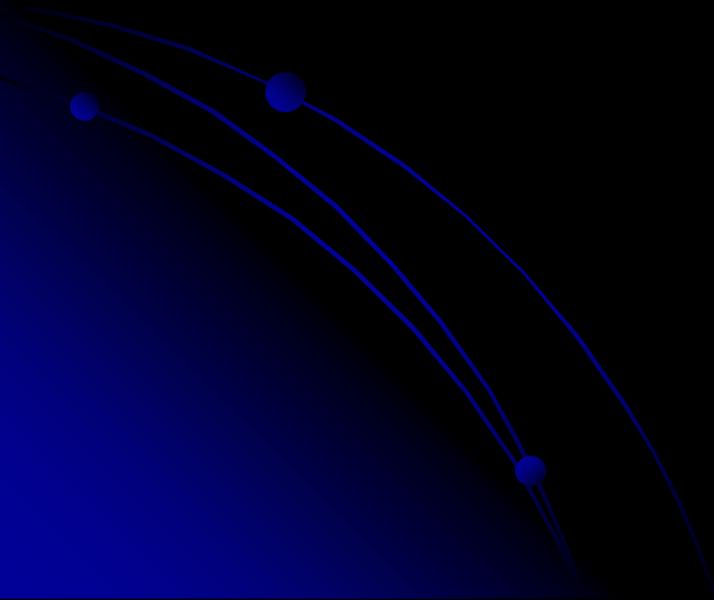
The 4 Forces

- 1) Strong Atomic Force- holds the nucleus of the atom together (likes repel)
- 2) Electromagnetic- the force between particles with charge/magnetism (opposites attract)
- 3) Weak Atomic Force - controls the radioactive decay of atomic nuclei
- 4) Gravitation – attraction/force between particles with mass

Have scientists combined any of these forces?

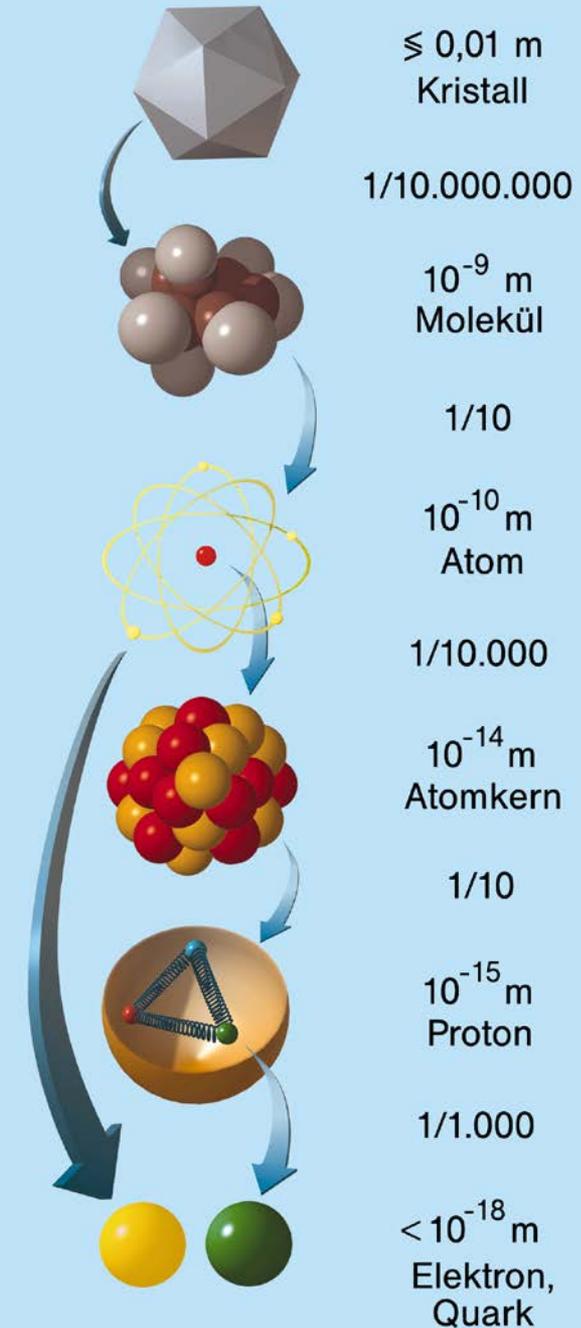
Yes!

We have combined electromagnetism & the weak nuclear force. This forms a z-particle, this it is called the electroweak force.



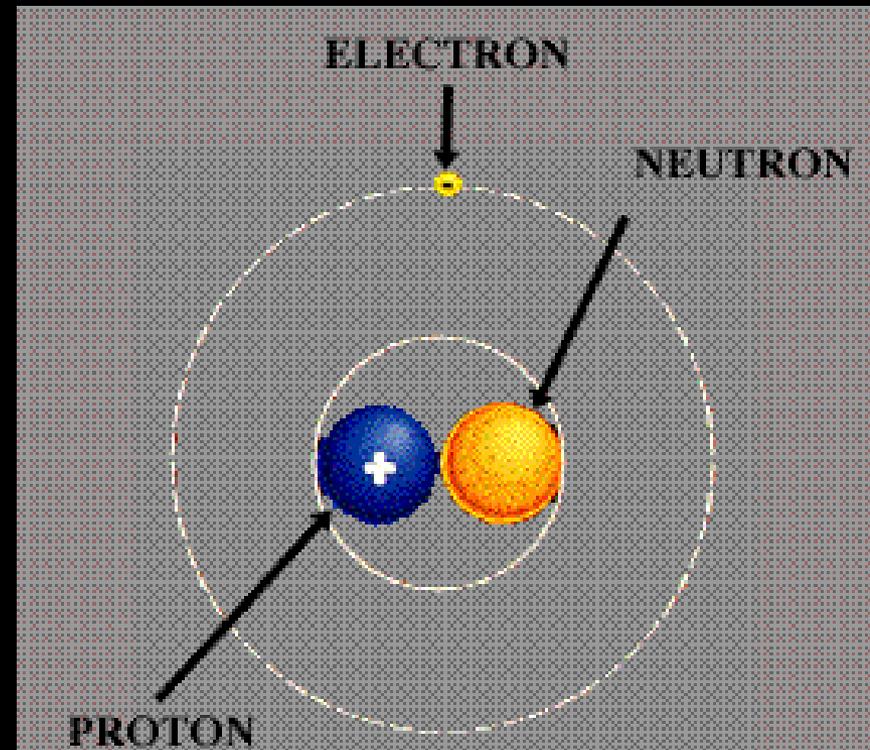
Step 3:

- Quarks were produced-
the most elementary
particles of atoms.



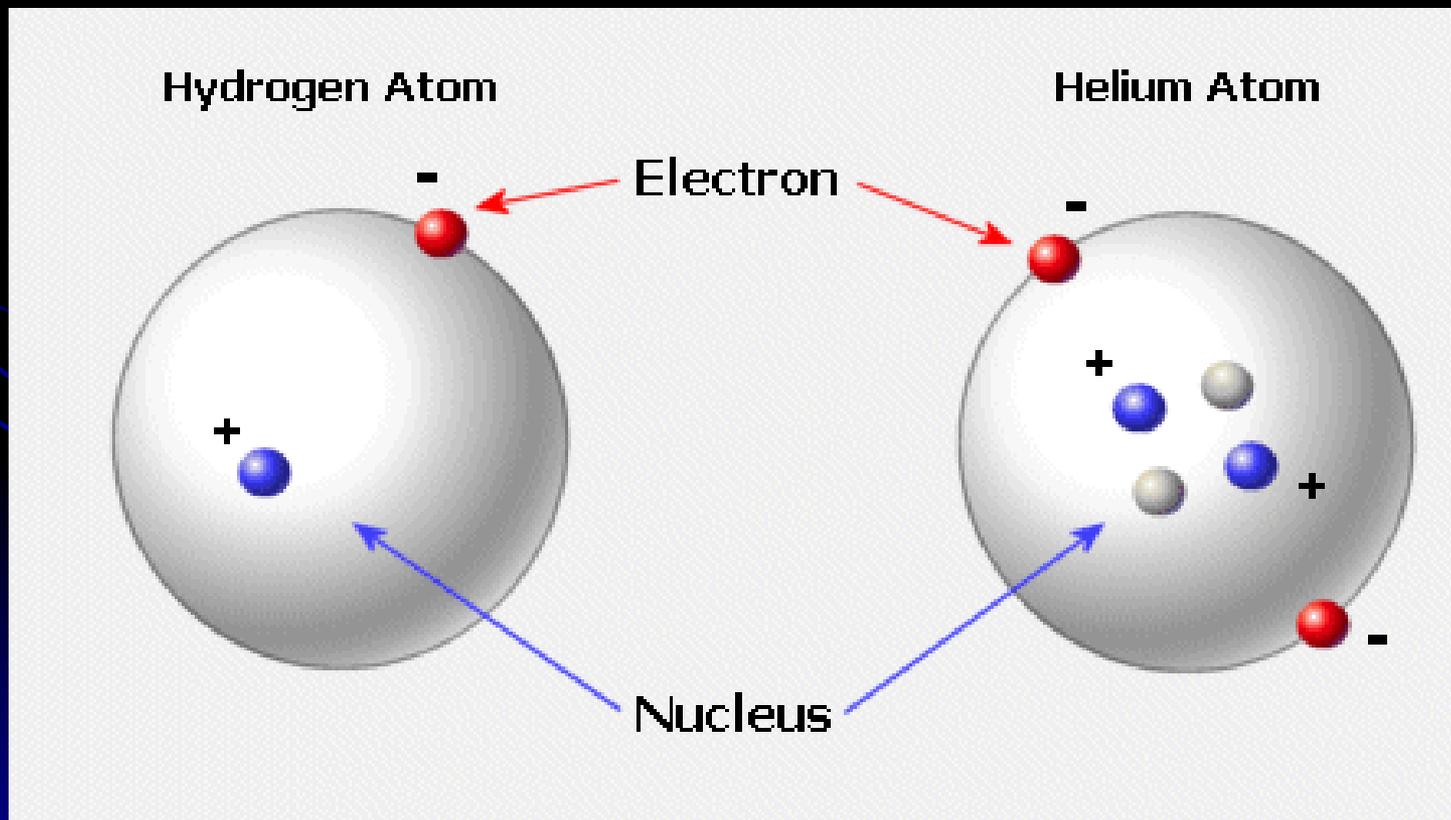
Step 4:

- Quarks combine to form protons (p), neutrons (n) & electrons (e⁻)



Step 5

- p^+ , n , & e^- combine & form Hydrogen (H) & Helium (He)



Step 6

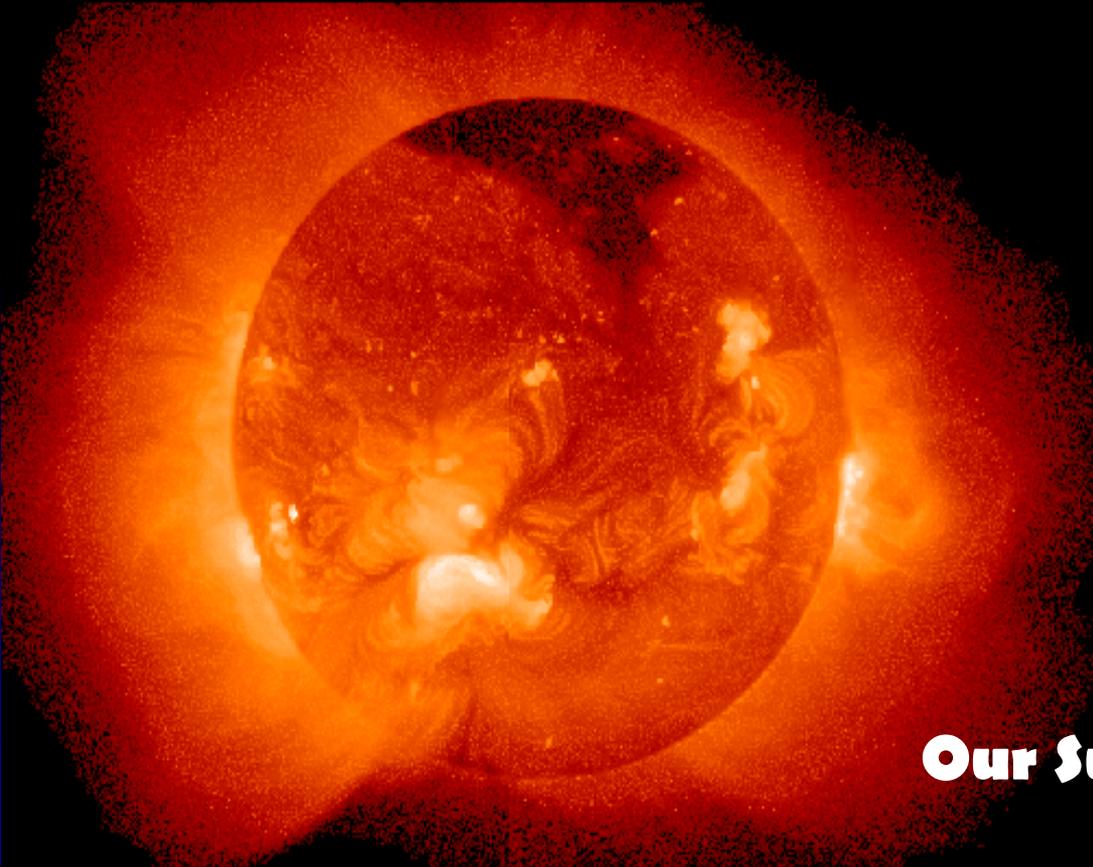
- H & He condense into gas clouds (nebulas)

The image shows the Eagle Nebula, a large interstellar cloud of dust and gas. It features several prominent, dark, pillar-like structures (proplydads) that rise from a glowing, blue-tinted nebula. The pillars are illuminated from the side, creating a dramatic silhouette effect. The background is a deep blue with scattered stars and faint nebular structures.

Eagle Nebula (gas cloud)
7,000 light-years from Earth
The left tower is 3 LY long.
One LY is 5,800,000,000 miles

Step 7

- The gas clouds evolve into stars.



Our Sun

Step 8

- Stars combine into billions & billions of galaxies!



A spiral galaxy



**This is the Andromeda Galaxy, our nearest large neighbor galaxy.
NASA**